

# Correction to: Observations of a Radio-Quiet Solar Preflare (vol 292, 2017)

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## Correction to: Observations of a Radio-Quiet Solar Preflare

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### Correction to: Solar Phys

<https://doi.org/10.1007/s11207-017-1175-3>

This article was published with an erroneous version of Figure 3. Please find in this document the correct version of the figure.

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Combined Radio and Space-based Solar Observations: From Techniques to New Results  
Guest Editors: Eduard Kontar and Alexander Nindos

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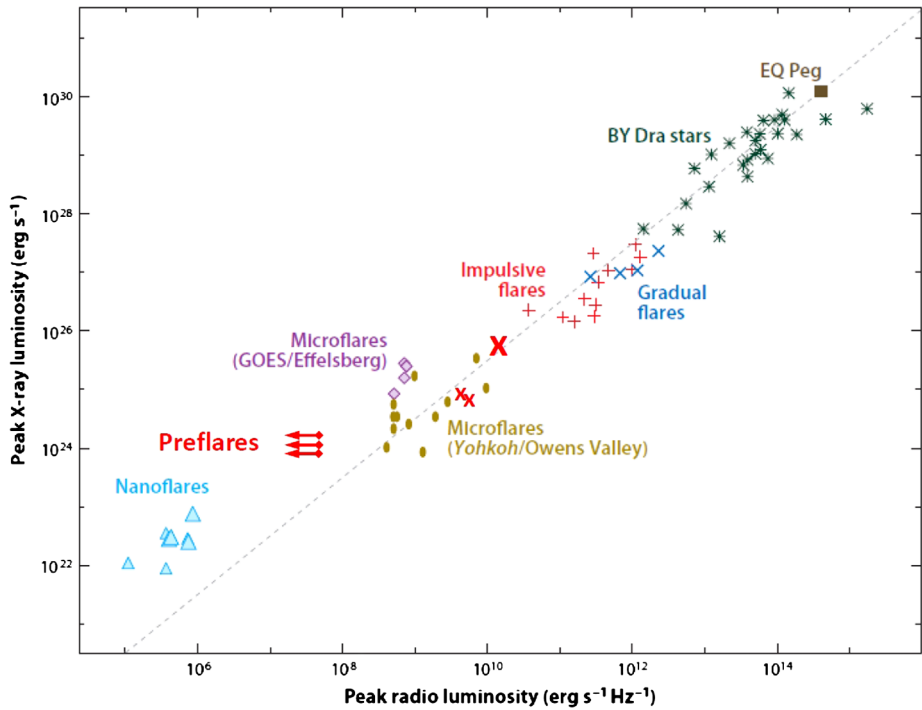
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**Figure 3** Comparison of the peak soft X-ray and 3.6 cm radio luminosity of solar and stellar flares. The upper limits derived from these preflare observations are shown by *red arrows*. Two peaks in the early main flare phase are indicated with a *lowercase red x*, the main peak with an *uppercase X*. We also show nanoflares in quiet regions (Krucker and Benz, 2000), microflares in active regions (Benz *et al.*, 1981; Fürst, Benz, and Hirth, 1982; Gary, Hartl, and Shimizu, 1997), impulsive and gradual flares (Benz and Güdel, 1994), quiescent BY Dra, dMe, and dKe stars (Güdel *et al.*, 1993), and a long-duration flare on the dMe star EQ Peg (Kundu *et al.*, 1988). The *dashed line* presents the relation given in Equation 1. (Graphics adapted from Benz and Güdel, 2010, and based on Krucker and Benz, 2000.)